Physics 480/581

Problem Session No. 4

Monday, 24 September, 2018

1. On a vector space of 1-forms over a flat manifold with coordinates $\{x, y, z, t\}$, choose as a basis the following:

$$\theta^1 \equiv dx + i \, dy, \quad \theta^2 \equiv dx - i \, dy, \quad \theta^3 \equiv dz - dt, \quad \theta^4 \equiv dz + dt.$$

For 2 timelike-separated events, A and B on the manifold, with coordinates $\{x_A^{\mu}\}$ and $\{x_B^{\nu}\}$ the (Lorentz invariant) interval between them would be given by

$$\mathbf{g} \equiv \boldsymbol{\eta}_{\mu\nu} \left(x_B^{\mu} - x_A^{\mu} \right) \left(x_B^{\nu} - x_A^{\nu} \right) \equiv -(\Delta \tau_{AB})^2 ,$$

assuming that τ is the parameter along a worldline of some observer who passed between A and B. An infinitesimal version of the interval, using $\{dx^{\mu}\}$ as a choice for a basis of 1-forms or the choice given above $\{\theta^{\alpha}\}$, would just be

$$\eta_{\mu\nu} dx^{\mu} dx^{\nu} \equiv Q_{\alpha\beta} \varrho^{\alpha} \varrho^{\beta} ,$$

where $Q_{\alpha\beta}$ is the form of what is usually referred to as the metric tensor relative to the first choice of basis 1-forms, as given in the beginning of this problem. Please determine the matrix which has the $Q_{\alpha\beta}$ as its elements.

2. Using the connection 1-forms for the Schwarzschild metric, in the orthonormal basis

the connection 1-forms are the following:

write out the 4 differential equations that define a geodesic in this spacetime, remembering that the connection 1-forms are skew-symmetric in their indices.

3. The 4-velocity can be written as

$$\frac{d}{d\tau} = \widetilde{u} = \frac{dx^{\mu}}{d\tau} \frac{\partial}{\partial x^{\mu}} .$$

If the coordinates $\{x^{\mu}\}$ are the usual spherical coordinates, $\{r, \theta, \varphi, t\}$, what is a matrix representation of the components of the 4-velocity?

4. In the orthonormal basis, show that the components of the 4-velocity are such that

$$\gamma = \mathcal{H} \frac{dt}{d\tau} \,, \quad v^{\hat{r}} = \frac{dr/dt}{\mathcal{H}} \,, \quad v^{\hat{\theta}} = \frac{r}{\mathcal{H}} \frac{d\theta}{dt} \,, \quad v^{\hat{\varphi}} = \frac{r \sin \theta}{\mathcal{H}} \frac{d\varphi}{dt} \,.$$

Remember that $(\widetilde{u})^2 = -1$.